

Energy Quantisation in Bulk Bouncing Tachyon

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Abstract

We argue that the closed string energy in the bulk bouncing tachyon background is to be quantised in a simple manner as if strings were trapped in a finite time interval. We discuss it from three different viewpoints; (1) the timelike continuation of the sinh-Gordon model, (2) the dual matrix model description of the $(1+1)$ -dimensional string theory with the bulk bouncing tachyon condensate, (3) the $c_L = 1$ limit of the timelike Liouville theory with the dual Liouville potential turned on. There appears to be a parallel between the bulk bouncing tachyon and the full S-brane of D-brane decay. We find the critical value λ_c of the bulk bouncing tachyon coupling which is analogous to $\lambda_o = \frac{1}{2}$ of the full S-brane coupling, at which the system is thought to be at the bottom of the tachyon potential.

1 Introduction

In recent years there has been a good amount of progress in understanding the dynamics of tachyons localised in space – the open string tachyon of D-brane decay [1] and twisted closed string tachyon of orbifold decay [2] (see also [3] for a review).

There is little understanding of what would happen when the bulk tachyon condenses. The bulk tachyon indicates the instability of the flat spacetime, so it is hard to imagine what kind of spacetime we would end up with after the decay. A common speculation is that the spacetime disappears as the bulk tachyon condenses. As closed strings propagate in the tachyon media in spacetime, they would feel harder to go through it than in the spacetime

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without tachyon condensate. If the endpoint of the tachyon decay is at $T_{blk} = \infty$, it would be too hard for strings to propagate at all in the infinitely stiff tachyon media in spacetime. In fact it may be noted that the dual matrix model description [4, 5] of the bulk tachyon condensate in $(1+1)$ -dimensions seems to suggest this speculation to be true, even though in $(1+1)$ -dimensions the closed string tachyon, being massless, does not indicate an instability and its condensation does not represent the decay process.

A class of marginal deformations of conformal field theory (CFT) can describe the tachyon condensation with certain profiles. However, these deformations are not a small perturbation, and in general it is not easy to solve the interacting CFTs with these deformations. So finding such exactly solvable CFTs can potentially lead to progress in understanding the tachyon dynamics. It was indeed the case for the open string tachyon of D-brane decay. A nontrivial exact $c = 1$ boundary CFT (BCFT) – the massless limit of the boundary sine-Gordon model – was solved in [6, 7]. Later it was applied by means of a timelike continuation to the bouncing open string tachyon (full S-brane) of D-brane decay by Sen [8, 9]. Subsequently a similar BCFT was studied in [10, 11] in its application to the rolling open string tachyon (half S-brane).¹ The BCFT descriptions of the full and half S-branes led to significant progress in understanding the D-brane decay.

In the bulk case, a nontrivial $c = 1$ CFT was found by Runkel and Watts [14] as a limit of the minimal CFTs. It was later proved by Schomerus and Fredenhagen [15, 16] that this CFT is indeed a $c = 1$ limit of the Liouville CFT [17, 18]. In the same spirit as Sen, it was then applied by means of a timelike continuation to the bulk rolling tachyon by Strominger and Takayanagi [19], and Schomerus [15]. These developments can potentially be important steps towards better understanding the bulk tachyon condensation.

In this paper we will consider the bulk bouncing tachyon from three different viewpoints; (1) the timelike continuation of the sinh-Gordon (ShG) model, (2) the dual matrix model description of the $(1+1)$ -dimensional string theory with the bulk bouncing tachyon condensate, (3) the $c_L = 1$ limit of the timelike Liouville theory with the dual Liouville potential turned on. We will study in particular the closed string spectrum in the bulk bouncing tachyon background, and propose that the closed string energy is to be quantised as if strings were trapped in a finite time interval.

The organisation of the paper is as follows. In section 2, we briefly remind the reader what the bulk bouncing tachyon is and its relation to the Liouville theory. In section 3, we study the timelike continuation of the ShG model and argue that the energy is to be quantised, based on Zamolodchikovs' argument of the momentum quantisation in the spacelike ShG model. In section 4, we study the dual matrix model of the $(1+1)$ -dimensional string theory with the bulk bouncing tachyon condensate. In particular, we discuss the spectrum of the theory and find an agreement with the result in section 3. In section 5, we make an attempt to construct the $c_L = 1$ limit of the Liouville theory with the dual Liouville potential turned

¹Following [12], we refer to the bouncing and rolling open string tachyon as full S-brane and half S-brane respectively – a coinage after S(pacelike)-brane of [13].

on. We discuss how the restriction on the energy spectrum can arise and how the results in the previous sections could potentially be consistent with this Liouville CFT. We summarise our results in section 6 and close the section with brief discussions. In appendix A, we show some of the details of the calculation in section 5.

2 Bulk bouncing tachyon

We will consider the Liouville theory with the dual Liouville potential turned on:

$$S = \frac{1}{4\pi} \int_{\Sigma} d^2x \sqrt{g} (g^{ab} \partial_a \phi \partial_b \phi + Q R_g \phi + 4\pi \mu e^{2b\phi} + 4\pi \tilde{\mu} e^{2\phi/b}) , \quad (2.1)$$

where $Q = b + 1/b$ and the central charge of this theory is $c_L = 1 + 6Q^2$.²

In particular, we are interested in its application to the bouncing bulk tachyon. So we will eventually focus on the case $b = i$ together with the Wick-rotation $\phi \rightarrow -i\phi$. This is a generalisation of the $c_L = 1$ timelike Liouville theory of Strominger and Takayanagi [19], and Schomerus [15]. In this case, the potential $V(-i\phi)$ becomes hyperbolic cosine, as one can see it by shifting the Wick-rotated Liouville field ϕ to $\phi + \frac{1}{4} \log(\tilde{\mu}/\mu)$, which gives

$$V_{b=i}(-i\phi) = \text{sign}(\mu) \times 2\lambda \cosh(2\phi) , \quad (2.2)$$

where $\lambda = \sqrt{\mu\tilde{\mu}}$ and we assume that $\mu\tilde{\mu} > 0$. Thus the $c_L = 1$ theory we would like to consider is a massless limit of the timelike sinh-Gordon (TShG) model.

In application to string theory, this corresponds to having the bulk tachyon profile to be hyperbolic cosine of the time coordinate ϕ . Classically the bulk tachyon T_{blk} comes in from $T_{blk} = \text{sign}(\mu) \times \infty$ in the infinite past ($\phi = -\infty$), reaches at the value of $T_{blk} = \text{sign}(\mu)\lambda$ at $\phi = 0$, and bounces back to $T_{blk} = \text{sign}(\mu) \times \infty$ in the infinite future ($\phi = +\infty$).

Without knowing the shape of the bulk tachyon potential, it may not be clear what this tachyon profile really represents. However, it is tempting to interpret it as a process of spacetime decay analogous to Sen's bouncing open string tachyon [8, 9] of D-brane decay. Thus in the critical string theory application, the bulk bouncing tachyon would be thought of as a process of homogeneous decay of the flat spacetime. We note that the analogy can be seen more manifestly in the dual matrix model description of the $(1+1)$ -dimensional string theory with the bulk bouncing tachyon condensate [4, 5], as we will review in section 4.

The main purpose of our paper is to understand a kinematical aspect – the spectrum – of the bulk bouncing tachyon background described by (2.1).³

²We will call μ cosmological constant (CC) and $\tilde{\mu}$ dual CC as commonly referred to as in literatures.

³The open string spectrum in the boundary massless sine-Gordon model – the spacelike continuation-back of the full S-brane – was computed in [7, 20].

3 Timelike sinh-Gordon model

Since the bulk bouncing tachyon is described by a massless limit of the TShG model, it would be a good starting point to look at the spacelike ShG model. On this score, we would like to put an emphasis on Zamolodchikovs' momentum quantisation in the ShG model (see Sect. 6 of Ref. [18] and also [21, 22, 23] for its application to other integrable models).

Consider the ShG model on the cylinder of circumference ϵ :

$$S_{ShG} = \frac{1}{4\pi} \int dt \int_0^\epsilon dx [(\partial_t \phi)^2 - (\partial_x \phi)^2 - 8\pi \lambda_{ShG} \cosh(2b\phi)] , \quad (3.1)$$

where $\phi(t, x) = \phi(t, x + \epsilon)$. This is a massive theory and the coupling $\lambda_{ShG}(> 0)$ has the dimension $\sim [\text{mass}]^{2+2b^2}$. It is convenient to apply the scaling transformation, $(t, x) \rightarrow \frac{2\pi}{\epsilon}(t, x)$. Since the scaling dimensions of ϕ and $\cosh(2b\phi)$ are 0 and $-2b^2$ respectively, the ShG action then becomes

$$S_{ShG} = \frac{1}{4\pi} \int dt \int_0^{2\pi} dx \left[(\partial_t \phi)^2 - (\partial_x \phi)^2 - 8\pi \lambda_{ShG} \left(\frac{\epsilon}{2\pi} \right)^{2+2b^2} \cosh(2b\phi) \right] . \quad (3.2)$$

First of all, the ShG potential is a confining potential, and thus one may expect to have the momentum to be quantised.

Zamolodchikovs made it more precise; We can regard the ShG model as the Liouville theory with the potential $\lambda_{ShG} e^{-2b\phi}$ on the left (in the ϕ -coordinate) and $\lambda_{ShG} e^{2b\phi}$ on the right. In fact the locations of the walls of the Liouville potentials are given by $\phi_{left} = \frac{1}{2b} \ln \left(\lambda_{ShG} (\epsilon/2\pi)^{2+2b^2} \right)$ on the left and $\phi_{right} = -\frac{1}{2b} \ln \left(\lambda_{ShG} (\epsilon/2\pi)^{2+2b^2} \right)$ on the right. As we go to the UV ($\epsilon \ll 1$), two walls are well separated, and in the middle ($\phi_{left} \ll \phi \ll \phi_{right}$) the potential is free. So if we send a free wave from the middle, say, to the right wall, it will be reflected as if in the Liouville theory with the potential $\lambda_{ShG} \phi^{2b\phi}$, then propagate through leftwards freely in the middle, be again reflected at the left wall this time as if in the Liouville theory with the potential $\lambda_{ShG} \phi^{-2b\phi}$, and then come back to the middle.

The wavefunction after the first reflection is given by

$$\Psi_P[\phi] \sim e^{2iP\phi} + (\epsilon/2\pi)^{-4iPQ} S_b(P) e^{-2iP\phi} , \quad (3.3)$$

where $S_b(P)$ is the Liouville reflection amplitude

$$S_b(P) = - \left(\pi \lambda_{ShG} \gamma(b^2) \right)^{-2iP/b} \frac{\Gamma(1 + 2iP/b) \Gamma(1 + 2iPb)}{\Gamma(1 - 2iP/b) \Gamma(1 - 2iPb)} . \quad (3.4)$$

Then the wavefunction after the second reflection follows

$$\Psi_P[\phi] \sim (\epsilon/2\pi)^{-8iPQ} S_{-b}(-P) S_b(P) e^{2iP\phi} + (\epsilon/2\pi)^{-4iPQ} S_b(P) e^{-2iP\phi} . \quad (3.5)$$

These two wavefunctions have to match up, which requires the consistency condition

$$(\epsilon/2\pi)^{-8iPQ} S_{-b}(-P) S_b(P) = (\epsilon/2\pi)^{-8iPQ} S_b(P)^2 = 1 . \quad (3.6)$$

This is the momentum quantisation.

Now we can repeat a similar analysis for the TShG model obtained by $b \rightarrow i\beta$ and $\phi \rightarrow -i\phi$,

$$S_{TShG} = \frac{1}{4\pi} \int dt \int_0^{2\pi} dx \left[-(\partial_t \phi)^2 + (\partial_x \phi)^2 - 8\pi \lambda_{TShG} \left(\frac{\epsilon}{2\pi} \right)^{2-2\beta^2} \cosh(2\beta\phi) \right] , \quad (3.7)$$

but with $\lambda_{TShG} < 0$. In order to have the confining potential, λ_{TShG} must be negative, since the sign of the kinetic term is reversed for the timelike theory.

The energy is negative definite, and the action of the timelike theory, after the Wick-rotation of the world-sheet time, is bounded from above. So we need to find, if possible, a good prescription to define this timelike theory. Fortunately, the timelike Liouville theories which are ingredients of the TShG model (3.7) were constructed by Schomerus [15], Strominger and Takayanagi [19]. They assumed that $\lambda_{TShG} > 0$, but there appears to be no obstruction to continue λ_{TShG} to negative value in their results. Furthermore, the spacelike theory from which the timelike theory is analytically continued has an oscillating potential.

As far as the reflection amplitude is concerned, it can be obtained in a straightforward manner, simply replacing b by $i\beta$ together with the analytic continuation $P \rightarrow -i\omega$ in (3.4). Then the argument which leads to the momentum quantisation is exactly the same as in the spacelike case. Thus we have

$$(\epsilon/2\pi)^{-8i\omega Q_\beta} S_{-i\beta}(i\omega) S_{i\beta}(-i\omega) = (\epsilon/2\pi)^{-8i\omega Q_\beta} S_{i\beta}(-i\omega)^2 = 1 , \quad (3.8)$$

where $Q_\beta = \beta - 1/\beta$. This is the energy quantisation we would like to claim.

In the massless limit $\beta \rightarrow 1$, this becomes⁴

$$\lim_{\beta \rightarrow 1} (\pi \lambda_{TShG} \gamma(-\beta^2))^{4i\omega} = 1 . \quad (3.9)$$

Since the $\lim_{\beta \rightarrow 1} \gamma(-\beta^2)$ diverges, we need to take the $\lambda_{TShG} \rightarrow 0$ limit such that the $\lim_{\beta \rightarrow 1} (\pi \lambda_{TShG} \gamma(-\beta^2))$ is kept finite. We will give more precise prescription to take this limit below, but let us denote $\lim_{\beta \rightarrow 1} (\pi \lambda_{TShG} \gamma(-\beta^2))$ as $(\lambda_{TShG})_{ren}$ for the moment. Then the energy quantisation (3.9) reads

$$\omega = \frac{\pi n}{2 \ln(\lambda_{TShG})_{ren}} , \quad n \in \mathbb{Z}_+ . \quad (3.10)$$

Since the positive and negative ω are related and thus not independent due to the reflection, we have taken only the positive integer n . The zero mode $n = 0$ corresponds to a constant wavefunction and so should be removed from a complete set of states in the Hilbert space with the discrete spectrum.

⁴More precisely, as we will see in section 5, the $\lim_{\beta \rightarrow 1} S_{i\beta}(i\omega)$ and $\lim_{\beta \rightarrow 1} S_{-i\beta}(i\omega)$ are given by the first and second equation in (5.34) respectively, where $\omega_1 = \omega_2 = \omega$, $\omega_3 = 0$ and $\lambda_{TShG} = \mu = \tilde{\mu}$. Thus the expression (3.9) is slightly heuristic.

In [24], Fredenhagen and Schomerus studied the minisuperspace approximation of the rolling and bouncing tachyons. In particular, in the bouncing tachyon case, they found the discrete energy spectrum. They considered the case of $\lambda_{TShG} > 0$. In that case, the potential is not confining, but one still needs to impose the boundary conditions at $\phi = \pm\infty$, since a classical particle in the exponential potentials can reach at $\phi = \pm\infty$ in a finite time. As the system is then effectively in a finite box, the energy is quantised. Our quantisation condition is different from theirs. This is not a contradiction, since the boundary conditions at $\phi = \pm\infty$ would not yield the same effect as the confining potential. Also the minisuperspace approximation is a good approximation in the semi-classical limit $\beta \rightarrow 0$, while we are considering the $\beta \rightarrow 1$ limit.

As long as the theory is massive, two walls are well separated in the UV ($\epsilon \ll 1$), which validates the above argument in the UV. Moreover, in this case, there is an alternative check of the momentum quantisation (3.6) by the Thermodynamic Bethe Ansatz for the spacelike theory.

In the massless limit, the theory does not depend on the scale ϵ . So the picture of the well-separated walls is only true when $|(\lambda_{TShG})_{ren}| \ll 1$. However, we would like to argue that the energy quantisation (3.10) is valid for generic value of $(\lambda_{TShG})_{ren}$. In the following sections, we will provide two alternative arguments which support our claim.

4 The matrix model

We would like to use the matrix model of $(1+1)$ -dimensional string theory in order to learn the $c_L = 1$ Liouville theory with the dual Liouville potential turned on. This is an idea explored in [4] where the $c = 1$ matrix model was used to study the Liouville theory with $c_L < 25$.

The $(1+1)$ -dimensional string theory with the linear dilaton is dual to the double scaled matrix quantum mechanics (MQM) [25, 26, 27, 28, 29] and [30, 31]. The $U(N)$ MQM is equivalent to the system of N non-interacting fermions trapped in the MQM potential (fermi liquid) [32]. In the double scaling limit [26, 27, 28, 29], N is taken to the infinity and the fermi level $-\mu_{bare}$ as measured from the top of the potential is taken to zero, while keeping $N\mu_{bare}(= 1/g_s)$ fixed finite. Then the double scaled MQM is the second quantised free fermion theory in the inverted harmonic oscillator potential:

$$H_F = \frac{1}{2} \int_{-\infty}^{\infty} dx \left(\partial_x \Psi^\dagger \partial_x \Psi - x^2 \Psi^\dagger \Psi \right) , \quad (4.1)$$

where the fermion satisfies the commutation relations

$$\{ \Psi^\dagger(x), \Psi(x') \} = \delta(x - x') . \quad (4.2)$$

The vacuum – the flat spacetime with the linear dilaton and the tachyon condensate $T = -\mu x e^{2x}$ – corresponds to the fermi sea, $h(x, p) \equiv \frac{1}{2}(p^2 - x^2) \leq -\mu$, or equivalently $-\sqrt{x^2 - 2\mu} \equiv p_- \leq p \leq p_+ \equiv \sqrt{x^2 - 2\mu}$.

It is possible to deform the fermi surface. To see it, let us consider the Hamilton-Jacobi equation of this system. It is given by

$$\partial_t S = -h(x, \partial_x S) = \frac{1}{2} (x^2 - (\partial_x S)^2) , \quad (4.3)$$

where $S(x; t)$ is Hamilton's principal function and $p = \partial_x S$. Taking the partial derivative of this equation w.r.t. x , we obtain the Euler equation [33]

$$\partial_t p = x - p \partial_x p . \quad (4.4)$$

The static solution $p = p(x)$ corresponds to the vacuum as we commented above, whereas the non-static solutions $p = p(x; t)$ deform the static fermi surface $p_{\pm} = \pm \sqrt{x^2 - 2\mu}$.

The bosonisation $\partial_x \phi = \pi \Psi^\dagger \Psi$ leads to Das-Jevicki-Sakita's collective field description of the double scaled MQM [34, 35]:

$$S_B = \int \frac{dt dx}{\pi} \left[\frac{(\partial_t \phi)^2}{2 \partial_x \phi} - \frac{1}{6} (\partial_x \phi)^3 + \left(\frac{1}{2} x^2 - \mu \right) \partial_x \phi \right] . \quad (4.5)$$

This is a $(1+1)$ -dimensional string field theory, and the solutions to the classical equation of motion,

$$\partial_t \left(\frac{\partial_t \phi}{\partial_x \phi} \right) - \frac{1}{2} \partial_x \left(\frac{\partial_t \phi}{\partial_x \phi} \right)^2 - \frac{1}{2} \partial_x (\partial_x \phi)^2 + x = 0 , \quad (4.6)$$

describe various different backgrounds. With the relations [33],

$$\partial_x \phi = \frac{p_+ - p_-}{2} , \quad \frac{\partial_t \phi}{\partial_x \phi} = -\frac{p_+ + p_-}{2} , \quad (4.7)$$

the bosonic equation of motion (4.6) is indeed equivalent to the fermionic ones

$$\partial_t p_{\pm} = x - p_{\pm} \partial_x p_{\pm} . \quad (4.8)$$

Hence the deformations of the fermi surface correspond to generating different backgrounds in the $(1+1)$ -dimensional string field theory [33].

Of our particular interest among non-static solutions to the Euler equation (4.8) is

$$(\sinh(t + t_0)x - \cosh(t + t_0)p) (\sinh(t - t_0)x - \cosh(t - t_0)p) = -\mu \sinh(2t_0) , \quad (4.9)$$

found in [4, 5], where $t_0 \geq 0$ is a constant (see also related works [36, 37, 38, 39, 40, 41, 42, 43, 44, 45] on other non-static backgrounds).⁵ Incidentally this solution can be generated from the static solution by a similarity transformation $(x, p) \mapsto U^{-1}(x, p)U$ where $U = e^{-\frac{\alpha}{2} w_2}$

⁵The non-static background (4.9) is an example of the spacetime with spacelike boundaries. Thus it offers an interesting laboratory for understanding the spacetime with cosmological horizon [44]. We would like to thank S. Das for bringing our attention to the paper [44].

with $t_0 = \frac{1}{4} \ln \tanh \alpha$ and $w_2 = w_{02} - w_{20}$ together with a scaling of μ [40]. The generators w_{rs} of the W_∞ algebra are given by

$$w_{rs} = \frac{1}{2} e^{(r-s)t} (x-p)^r (x+p)^s, \quad (4.10)$$

which satisfy

$$\{w_{rs}, w_{r's'}\}_{PB} = (rs' - sr') w_{r+r'-1, s+s'-1}. \quad (4.11)$$

We shall consider the bosonic string theory, so will restrict ourselves to one side of the fermi sea in the region $x < 0$. For large $-x$,⁶ it is easy to find from the relations (4.7) that

$$\partial_x \phi = -\mu x^{-1} + \frac{1}{2} (\tanh(t+t_0) - \tanh(t-t_0)) x + \mathcal{O}(x^{-3}). \quad (4.13)$$

The world-sheet tachyon T is related to the collective field ϕ , for $x^{-2}\phi \ll 1$, via the dictionary [33]

$$T + \frac{1}{2} = e^{2q}\phi + \mathcal{O}(e^{4q}), \quad (4.14)$$

up to the leg pole factor which is not relevant in our case. Here we have defined $x \equiv -e^{-q}$. Then we have [4, 5]

$$T = \mu q e^{2q} + \frac{1}{4} \left(\frac{\sinh(2t_0)}{\cosh(t+t_0) \cosh(t-t_0)} - 2 \right) \sim \mu q e^{2q} - \frac{1}{2} \lambda (e^{2t} + e^{-2t}), \quad (4.15)$$

where we have defined $\lambda_{MM} = e^{-2t_0} \leq 1$. The last expression is valid for $-|\ln \lambda_{MM}| \ll t \ll |\ln \lambda_{MM}|$ only for which $x^{-2}\phi \ll 1$ and the dictionary works. Thus the non-static solution (4.9) describes the bulk bouncing tachyon in the linear dilaton background [4, 5]. In particular, the last expression in (4.15) corresponds to the world-sheet theory of the direct product of two CFTs, the $c_L = 25$ ($b = 1$) Liouville and the $c_L = 1$ ($b = i$) limit of the Liouville theory (2.1) with the dual Liouville potential turned on.

The non-static solution (4.9) suggests the following interpretation of the bulk bouncing tachyon [4, 5]: We will set μ to zero for simplicity. But the following picture is essentially the same for finite μ . In the infinite past $t = -\infty$, the upper p_+ and lower half p_- of the fermi surface degenerate, that is, $p_\pm = -x$. This means that the fermi sea, $p_- \leq p \leq p_+$, is drained out. There is nothing to be excited from the empty fermi sea – no string propagates. Since the dictionary (4.14) breaks down in particular at $t = -\infty$, we do not know it for sure but it is tempting to consider that the tachyon is at $T = \infty$ at this point. Strings cannot propagate in the infinitely stiff tachyon media in spacetime. As time goes by, the fermi sea is being filled until $t = 0$ at which $p_\pm = \pm \tanh(t_0)x$ and it then starts draining. It keeps on draining and is again emptied out in the infinite future $t = +\infty$ at which $p_\pm = x$.

⁶More precisely,

$$-x \gg \sqrt{\frac{4\mu \cosh(t+t_0) \cosh(t-t_0)}{\sinh(2t_0)}}. \quad (4.12)$$

Rephrasing the process of the bulk bouncing tachyon, the spacetime (or universe) starts out from nothing in the infinite past, is then created and “expands” until $t = 0$, and then “contracts” until it eventually dies off to nothing in the infinite future. Note that there is a parallel between this picture and the full S-brane of D-brane decay. In the latter case, an unstable D-brane is created and then annihilated, instead of the spacetime (or universe) [12].

Note also that when $t_0 = 0$ thus $\lambda_{MM} = 1$, the upper and lower half of the fermi surface always degenerate, that is, $p_{\pm} = \tanh(t)x$. Thus the fermi sea is drained out at all time. This point $\lambda_{MM} = 1$ is similar to $\lambda_o = \frac{1}{2}$ for the full S-brane at which the system is thought to be at the bottom of the tachyon potential.

4.1 The spectrum

We would now like to understand the closed string spectrum in the background (4.9). We will thus expand the collective field action (4.5) about the background ϕ_0 as

$$\phi(t, x) = \phi_0(t, x) + \sqrt{\pi}\eta(t, x) , \quad (4.16)$$

and look at the quadratic part w.r.t. the fluctuation η . The quadratic action is given by

$$S_2 = \frac{1}{2} \int \frac{dtdx}{\partial_x \phi_0} \left[\left(\partial_t \eta - \frac{\partial_t \phi_0}{\partial_x \phi_0} \partial_x \eta \right)^2 - (\partial_x \phi_0)^2 (\partial_x \eta)^2 \right] . \quad (4.17)$$

Following [37], we introduce a new coordinate system (τ, y) by

$$y = \phi_0(t, x) \quad \text{and} \quad \tau = t . \quad (4.18)$$

Then the quadratic action simplifies to the diagonal form

$$S_2 = \frac{1}{2} \int d\tau dy \left[\frac{1}{(\partial_x \phi_0)^2} (\partial_\tau \eta)^2 - (\partial_x \phi_0)^2 (\partial_y \eta)^2 \right] . \quad (4.19)$$

With this form, the fluctuation spectrum can be found by solving the equation of motion

$$\left[\partial_\tau \left(\frac{1}{(\partial_x \phi_0)^2} \partial_\tau \right) - \partial_y ((\partial_x \phi_0)^2 \partial_y) \right] \eta(\tau, y) = 0 . \quad (4.20)$$

In our application, we have the explicit form

$$y \sim -\frac{\sinh(2t_0)}{4 \cosh(t+t_0) \cosh(t-t_0)} x^2 \quad \text{for} \quad , \quad -x \sqrt{\frac{\sinh(2t_0)}{4\mu \cosh(t+t_0) \cosh(t-t_0)}} \gg 1 . \quad (4.21)$$

One can then easily find that the equation of motion takes the form

$$[\partial_\chi^2 - \partial_{\ln(-y)}^2] \eta(\chi, \ln(-y)) \sim 0 , \quad \text{for} \quad -y \gg \mu , \quad (4.22)$$

where we have introduced a new time coordinate χ by

$$e^\chi = \frac{\cosh(\tau + t_0)}{\cosh(\tau - t_0)} \quad (4.23)$$

In fact the coordinate system $(\chi, \ln(-y))$ for $-y \gg \mu$ is an example of the flat coordinate in [37]. The quadratic action becomes that in the flat metric in terms of $(\chi, \ln(-y))$.

Note that since the r.h.s. of (4.23) increases monotonically as a function of τ , the new time coordinate χ has a finite range,

$$-2t_0 \leq \chi \leq 2t_0, \quad (4.24)$$

Note also that

$$\chi \sim 2\tau, \quad \text{for} \quad -t_0 \ll \tau \ll t_0. \quad (4.25)$$

The fluctuation equation of motion (4.22) reduces to

$$[\partial_t^2 - \partial_q^2] \eta(t, q) \sim 0 \quad \text{for} \quad -t_0 \ll t \ll t_0 \quad \text{and} \quad -q \gg 1, \quad (4.26)$$

where the dictionary (4.15) is valid. This is the region far away from the Liouville wall and the confining walls of the bouncing tachyon, where the theory is approximately free.

Since the time χ has the finite range, we need to impose the boundary conditions at $\chi = \pm 2t_0$. It seems natural to impose the Dirichlet boundary condition,

$$\eta(\chi = \pm 2t_0, \ln(-y)) = 0. \quad (4.27)$$

Then we have

$$\eta \propto e^{i\omega\chi} - e^{4it_0\omega} e^{-i\omega\chi} = e^{i\omega\chi} - \lambda_{MM}^{2i\omega} e^{-i\omega\chi}, \quad (4.28)$$

with the closed string energy ω being quantised by

$$\omega = \frac{\pi n}{4t_0} = \frac{\pi n}{2 \ln \lambda_{MM}}, \quad n \in \mathbb{Z}_+. \quad (4.29)$$

Since the free wave equation (4.22) takes the form (4.26) with the relation (4.25) in the weak field approximation, the energy ω here is the same as the one in the world-sheet theory (2.1). The same remark below (3.10) is applied regarding n being in \mathbb{Z}_+ .

Note first that the closed string wavefunction (4.28) precisely agrees with that of the massless limit of the TShG model, discussed in section 3, where the reflection amplitude is given by $-(\lambda_{TShG})_{ren}^{2i\omega}$. This agreement may suggest the Dirichlet boundary condition to be the right choice. Secondly we find the exact agreement between the energy quantisation (4.29) in the matrix model and the one (3.10) in the TShG model.

Incidentally the bulk rolling tachyon can be obtained by first shifting $\tau \rightarrow \tau + \tau_0$ and $\chi \rightarrow \chi + 2\tau_0$, and then taking the limit

$$\lambda_{MM} \rightarrow 0, \quad \tau_0 \rightarrow -\infty, \quad \text{keeping} \quad \kappa \equiv \lambda_{MM} e^{-2\tau_0} \quad \text{fixed finite}, \quad (4.30)$$

which corresponds to the time-dependent part of the tachyon condensate to be $T = -\frac{1}{2}\kappa e^{-2t}$. From (4.23), one can see that

$$\ln \kappa \leq \chi . \quad (4.31)$$

Thus the time direction is a half-line in the $(\chi, \ln(-y))$ coordinate. Imposing the Dirichlet boundary condition at $\chi = \ln \kappa$ yields

$$\eta \propto e^{i\omega\chi} - \kappa^{2i\omega} e^{-i\omega\chi} , \quad (4.32)$$

with $\omega \in \mathbb{R}_+$, which is in agreement with the reflection amplitude in the $c_L = 1$ timelike Liouville theory [19, 15].

The flat coordinate $(\chi, \ln(-y))$ is particularly a “good” frame at least for $-y \gg \mu$. To see it, let us look at the interaction part of the expansion of the action. One can find that it is given by

$$S_{int} \sim \int \frac{d\chi d\ln(-y)}{\pi} \left[-\frac{1}{6} e^{-\ln(-y)} (\sqrt{\pi} \partial_{\ln(-y)} \eta)^3 + \frac{1}{2} (\sqrt{\pi} \partial_{\chi} \eta)^2 \sum_{n=1}^{\infty} e^{-n \ln(-y)} (-\sqrt{\pi} \partial_{\ln(-y)} \eta)^n \right] . \quad (4.33)$$

Thus the interaction becomes time-independent, essentially the same as that of the static background, as in the case of [39] (see also a related discussion [42]). In particular the perturbative treatment seems to make sense as long as $-y \gg \mu$ even near $\chi \sim \pm 2t_0$ (thus $|t| \gg t_0$), which we believe validates our result.

Finally another point to make is that the flat coordinate $(\chi, \ln(-y))$ manifests the matrix cosmology picture of [5]. The spacetime is (conformally) flat, but the time interval in which the universe exists varies with the strength $\lambda_{MM} = e^{-2t_0}$ of the bulk bouncing tachyon as $-|\ln \lambda_{MM}| \leq \chi \leq |\ln \lambda_{MM}|$. At $\lambda_{MM} = 1$ the universe does not exist at all, which is analogous to $\lambda_o = \frac{1}{2}$ for the full S-brane of D-brane decay. In fact the energy quantisation (4.29) indicates that there is no excitation to be allowed at $\lambda_{MM} = 1$.

5 Timelike $c_L = 1$ Liouville

We now try to understand the energy quantisation directly from the construction of the $c_L = 1$ (thus $b = i$) limit of the timelike Liouville theory with the dual Liouville potential turned on. Our attempt will end incomplete. We will, however, be able to show how the restriction on the spectrum comes about and that the energy quantisation we found in the previous sections could be consistent with that restriction.

On this score, Teschner’s construction [46] of 2 and 3-pt functions is particularly useful. In short, this construction provides two different (dual) ways to obtain the 2 and 3-pt functions via the recursion relations. Two results then have to match up. This constraint could lead to the restriction on the spectrum. The condition from the 2-pt function consistency will be given in (5.15) and is in fact similar to the ones (3.6) and (3.8) of the (T)ShG model. In the case of $b = i$, it is very easy to see that (5.20) (the $b = i$ limit of (5.15)) could take exactly

of the same form as (3.9) if an appropriate choice of the function $f_{i/2}(\alpha|R)$ is made. We will now discuss it in more detail.

Teschner's construction is to consider auxiliary $(n+1)$ -pt functions with one insertion of the $(2,1)$ or $(1,2)$ degenerate field in order to calculate n -pt functions. The $(2,1)$ and $(1,2)$ degenerate fields are given by $\psi_{2,1} \equiv V_{-b/2}$ and $\psi_{1,2} \equiv V_{-1/2b}$ respectively, where we have defined

$$V_\alpha = e^{2\alpha\phi} , \quad (5.1)$$

with the conformal dimension $\Delta_\alpha = \alpha(Q - \alpha)$. The operator product expansions (OPEs) of any primary field with $\psi_{2,1}$ and $\psi_{1,2}$ involve only two conformal families [47],

$$V_\alpha V_{-b/2} = C_{\alpha,-b/2}^{\alpha-b/2} [V_{\alpha-b/2}] + C_{\alpha,-b/2}^{\alpha+b/2} [V_{\alpha+b/2}] , \quad (5.2)$$

$$V_\alpha V_{-1/2b} = C_{\alpha,-1/2b}^{\alpha-1/2b} [V_{\alpha-1/2b}] + C_{\alpha,-1/2b}^{\alpha+1/2b} [V_{\alpha+1/2b}] . \quad (5.3)$$

The structure constant $C_{\alpha,-b/2}^{\alpha+b/2}$ is given explicitly by [48]

$$C_{\alpha,-b/2}^{\alpha+b/2} = -\pi\mu \frac{\gamma(2b\alpha - 1 - b^2)}{\gamma(-b^2)\gamma(2b\alpha)} , \quad (5.4)$$

where $\gamma(x) = \Gamma(x)/\Gamma(1-x)$, and $C_{\alpha,-b/2}^{\alpha-b/2}$ and $C_{\alpha,-1/2b}^{\alpha-1/2b}$ are set to be unity by the choice of normalization.

A remark is in order; As we will show the details in Appendix A, the formula (5.4) for the structure constant is invalid at $b = i$. This is because there are an infinite number of contributions from the screening integrals at the point $b = i$, since the combination $\int d^2z e^{2b\phi(z)} \int d^2w e^{2\phi(w)/b}$ of two screening integrals has zero net charge. Summing them up then yields

$$C_{\alpha,-i/2}^{\alpha+i/2} = \lim_{b \rightarrow i} \frac{\pi\mu\gamma(b^2)}{(1 - (\pi\mu\gamma(b^2))(\pi\tilde{\mu}\gamma(1/b^2)))^2} . \quad (5.5)$$

The precise prescription of taking the $b \rightarrow i$ limit was given by Schomerus [15] (and Fredenhagen and Schomerus [16]) in their construction of 3-pt functions; setting $b^2 = -1 + i\epsilon$ and then sending $\epsilon \rightarrow 0_+$, in order to approach the limit from the well-defined regime of the 3-pt functions. Then we have

$$\begin{aligned} \lim_{b \rightarrow i} \gamma(b^2) &= \lim_{\epsilon \rightarrow 0_+} \gamma(-1 + i\epsilon) = \lim_{\epsilon \rightarrow 0_+} \left(\frac{i}{\epsilon} \right) , \\ \lim_{b \rightarrow i} \gamma(1/b^2) &= \lim_{\epsilon \rightarrow 0_+} \gamma(-1 - i\epsilon) = \lim_{\epsilon \rightarrow 0_+} \left(-\frac{i}{\epsilon} \right) . \end{aligned}$$

As we motivated ourselves in the TShG argument above, we would like to consider $\mu, \tilde{\mu} < 0$. In order to have a finite result, we renormalise the CCs, as in Strominger and Takayanagi [19], by setting $\mu = \mu_{ren}\epsilon$ and $\tilde{\mu} = \tilde{\mu}_{ren}\epsilon$. Then the structure constant in the $b \rightarrow i$ limit is given by

$$C_{\alpha,-i/2}^{\alpha+i/2} = \frac{i\pi\mu_{ren}}{(1 - (\pi\mu_{ren})(\pi\tilde{\mu}_{ren}))^2} , \quad (5.6)$$

where $\mu_{ren}, \tilde{\mu}_{ren} < 0$.

Similarly the dual structure constant $C_{\alpha, -1/2b}^{\alpha+1/2b}$ can be calculated as [48]

$$C_{\alpha, -1/2b}^{\alpha+1/2b} = -\pi\tilde{\mu} \frac{\gamma(2\alpha/b - 1 - 1/b^2)}{\gamma(-1/b^2)\gamma(2\alpha/b)} , \quad (5.7)$$

for $b \neq i$, and

$$C_{\alpha, i/2}^{\alpha-i/2} = \frac{-i\pi\tilde{\mu}_{ren}}{(1 - (\pi\mu_{ren})(\pi\tilde{\mu}_{ren}))^2} , \quad (5.8)$$

for $b = i$. They can be obtained by replacing b by $1/b$ and μ by $\tilde{\mu}$ in the expression of $C_{\alpha, -b/2}^{\alpha+b/2}$.

5.1 Two-Point Function

The two-point function in Liouville theory takes the form

$$|z|^{2\Delta_{\alpha_1}} \langle V_{\alpha_1}(z) V_{\alpha_2}(0) \rangle = \delta(\alpha_1 + \alpha_2 - Q) + D(\alpha_1) \delta(\alpha_1 - \alpha_2) . \quad (5.9)$$

The first term corresponds to $|z|^{2\Delta_{\alpha}} \langle V_{\alpha}(z) V_{Q-\alpha}(0) \rangle = 1$ which defines the normalisation, $C_{\alpha, -b/2}^{\alpha-b/2} = C_{\alpha, -1/2b}^{\alpha-1/2b} = 1$, that we adopted above.

We now consider an auxiliary 3-pt function with an insertion of the (2,1) degenerate field,

$$G_3(z_1, z_2, z; \alpha_1, \alpha_2) \equiv \langle V_{\alpha_1}(z_1) V_{\alpha_2+b/2}(z_2) V_{-b/2}(z) \rangle . \quad (5.10)$$

The associativity and locality require that two ways of taking the OPEs (1) $z \rightarrow z_1$ and (2) $z \rightarrow z_2$ in (5.10) give the same result, which in this case yields [48]

$$D(\alpha) = C_{\alpha, -b/2}^{\alpha+b/2} D(\alpha + b/2) . \quad (5.11)$$

This can be solved to

$$D(\alpha) = f_{b/2}(\alpha|R) (\pi\mu\gamma(b^2))^{(Q-2\alpha)/b} \frac{\gamma(2b\alpha - b^2)}{b^2\gamma(2 - 2\alpha/b + 1/b^2)} , \quad (5.12)$$

where the function $f_{\nu}(\alpha|R)$ is a two-parameter (ν and R) family of periodic functions of α with period ν . The parameter R is defined by $R = \ln \left[(\pi\tilde{\mu}\gamma(1/b^2))^b / (\pi\mu\gamma(b^2))^{1/b} \right]$ and is an “interval” between two Liouville walls.⁷ By a scaling argument, the periodic function $f_{\nu}(\alpha|R)$ can depend on μ and $\tilde{\mu}$ only through R . In the $b = i$ case, as we discussed above, we have to deal with it separately. However the result is as simple as replacing $\pi\mu\gamma(b^2)$ by (5.6) and otherwise setting $b = i$ in this expression.

A similar consideration by using the (1,2) degenerate field $V_{-1/2b}$ yields

$$D(\alpha) = g_{1/2b}(\alpha|R) (\pi\tilde{\mu}\gamma(1/b^2))^{b(Q-2\alpha)} \frac{b^2\gamma(2\alpha/b - 1/b^2)}{\gamma(2 - 2b\alpha + b^2)} , \quad (5.13)$$

⁷This is because formally the Liouville wall is defined by $\mu e^{2b\phi_0} \sim 1$, and the dual Liouville wall by $\tilde{\mu} e^{2\phi_0/b} \sim 1$.

The function $g_\nu(\alpha|R)$ is defined similarly as $f_\nu(\alpha|R)$. Again, in the $b = i$ case, we replace $\pi\tilde{\mu}\gamma(1/b^2)$ by (5.8).

The periodic functions f_ν and g_ν are not independent but related by the \mathbb{Z}_2 -symmetry in (2.1) under $b \leftrightarrow 1/b$ together with $\mu \leftrightarrow \tilde{\mu}$. It is easy to show, for instance, by a heuristic manipulation of the path integral expression of $D(\alpha)$, that

$$f_{1/2b}(\alpha|R) = g_{1/2b}(\alpha|-R) . \quad (5.14)$$

We have obtained the 2-pt function $D(\alpha)$ in two different ways, one (5.12) by using the (2,1) degenerate field and the other (5.13) by the (1,2) degenerate field. They must give the same result. The second γ -function factors are shown to be the same. Thus we have the constraint

$$f_{b/2}(\alpha|R) = e^{R(Q-2\alpha)} f_{1/2b}(\alpha|-R) . \quad (5.15)$$

Note also that there is another \mathbb{Z}_2 -symmetry under $b \leftrightarrow -b$ together with $\alpha \leftrightarrow -\alpha$ and $\phi \leftrightarrow -\phi$, which leads to a constraint

$$f_{-\nu}(-\alpha|-R) = f_\nu(\alpha|R) , \quad (5.16)$$

where $\nu = b/2$ or $1/2b$. This can be used when we discuss the $b = i$ case later.

There are two ways to think about the consistency condition (5.15); (1) it can be regarded as a condition on the periodic functions $f_{b/2}(\alpha|R)$ and $f_{1/2b}(\alpha|-R)$, should the (5.15) hold for arbitrary momentum α . (2) if we are to give up the first possibility, it may be thought of as the restriction on the allowed momentum α for given periodic functions $f_{b/2}(\alpha|R)$ and $f_{1/2b}(\alpha|-R)$.

In the first case, one can see that the eq.(5.15) implies

$$f_{b/2}(\alpha + 1/2b|R) = e^{-R/b} f_{b/2}(\alpha|R) , \quad (5.17)$$

$$f_{1/2b}(\alpha + b/2|-R) = e^{bR} f_{1/2b}(\alpha|-R) . \quad (5.18)$$

Thus $f_{b/2}(\alpha|R)$ and $f_{1/2b}(\alpha|-R)$ must be quasi-doubly periodic. When the ratio of two periods is not real, there actually exist an infinite number of such functions which can be constructed from the Jacobi Theta functions. The ratio, however, is b^2 , and thus this being not real corresponds to the strongly coupled gravity regime $1 < c_L < 25$. It is a very interesting problem to understand this regime, but we will not go into this case in the present paper.

In the case of $R = 0$, two periodic functions $f_{b/2}(\alpha|R)$ and $f_{1/2b}(\alpha|-R)$ are doubly periodic, and indeed the same from (5.15). Then it is well-known that, when the ratio of two periods b^2 is real and irrational, only such a function is constant. As noted in [46], although b^2 is rational for the much studied case of two-dimensional quantum gravity coupling to the minimal models, it still provides a supportive argument – if one assumes the continuity w.r.t. b – for uniqueness of the solution proposed in [17, 18].

We would like to make a comment here. Teschner's construction, as one can see from the above discussion, requires the presence of the dual Liouville potential. It may be interpreted

as that the dual Liouville potential is generated quantum mechanically [49]. The standard construction [18] of the Liouville theory further requires the dual CC to be at the self-dual point $R = 0$. This may be a manifestation of the $b \leftrightarrow 1/b$ duality of the 2 and 3-pt functions of [17, 18], as noted in [48]. However, the canonical quantisation [50] of the Liouville theory suggests that the Liouville field satisfies quantum mechanically the equation of motion without the dual Liouville potential. Thus it seems not too clear whether or not the standard Liouville theory should be thought of as the model with the dual Liouville potential tuned at the self-dual point.

5.1.1 The $b = i$ case

Since our primary interest is in the $b = i$ case, we shall restrict ourselves to this case. Two periods $b/2$ and $1/2b$ become minus of each other when $b = i$, so they are degenerate. If the constraints (5.17) and (5.18) were to be satisfied for arbitrary α , it would impose the factor e^{iR} to be 1. However it is impossible except for $R = 0$, since R is pure imaginary, if one assumes μ and $\tilde{\mu}$ to be real that is the case of physical interest.

Thus for generic R the only possibility is to regard the consistency condition (5.15) as the restriction on the momentum α .

The constraints (5.16) from the \mathbb{Z}_2 -symmetry, under $b \leftrightarrow -b$ together with $\alpha \leftrightarrow -\alpha$ and $\phi \leftrightarrow -\phi$, then relates two periodic functions by

$$f_{-i/2}(\alpha|-R) = f_{i/2}(-\alpha|R) . \quad (5.19)$$

Thus the restriction (5.15) on the momentum α now becomes

$$\frac{f_{i/2}(\alpha|R)}{f_{i/2}(-\alpha|R)} = e^{-4i\alpha \ln \lambda_{lft}} , \quad (5.20)$$

where

$$\ln \lambda_{lft} \equiv -\frac{iR}{2} = \ln \left(\frac{\lambda_{ren}}{(1 - \lambda_{ren}^2)^2} \right) , \quad \text{and} \quad \lambda_{ren} \equiv \pi \sqrt{\mu_{ren} \tilde{\mu}_{ren}} . \quad (5.21)$$

Notice its similarity to (3.9). It is clear that if $f_{i/2}(\alpha|R)$ is an even function, the constraint (5.20) is exactly of the same form as (3.9).

We will restrict ourselves to the case $\omega \equiv \alpha \in \mathbb{R}$ which corresponds to the timelike theory. Then the constraint (5.20) can be rewritten as

$$4\omega \ln \lambda_{lft} = 2\pi n + i \ln \left(\frac{f_{i/2}(\omega|R)}{f_{i/2}(-\omega|R)} \right) , \quad n \in \mathbb{Z} . \quad (5.22)$$

From the CFT viewpoint, the choice of the periodic function $f_{i/2}(\alpha|R)$, however, is rather arbitrary at this stage. If the function $f_{i/2}(\omega|R)$ is even or (quasi-)doubly periodic, that is, $f_{i/2}(\omega|R) = e^{2\pi i k} f_{i/2}(\omega + \pi/4 \ln \lambda_{lft}|R)$, where $k \in \mathbb{Z}$ and $f_{i/2}(0|R) < \infty$, and $f_{i/2}(\omega + i/2|R) = f_{i/2}(\omega|R)$, the constraint (5.22) reduces to the energy quantisation (3.10) and (4.29). We also note that the TShG and matrix model result (4.28) may suggest $f_{i/2}(\omega|R)$ to be simply 1. Then by the field renormalisation $V_\omega \rightarrow (e^{\pi i \tilde{\mu}_{ren}/\mu_{ren}})^{i\omega/2} V_\omega$ the reflection amplitude becomes $D(\omega) = -\lambda_{lft}^{2i\omega}$ in agreement with the results in the previous sections.

5.2 Three-Point Function

Next we repeat a similar analysis for the 3-pt functions. In this case, we consider an auxiliary 4-pt function with an insertion of the (2,1) degenerate field,

$$G_4(z_1, z_2, z_3, z; \alpha_1, \alpha_2, \alpha_3) \equiv \langle V_{\alpha_1}(z_1) V_{\alpha_2}(z_2) V_{\alpha_3}(z_3) V_{-b/2}(z) \rangle . \quad (5.23)$$

We will leave the details to Teschner's original paper [46]. The crossing symmetry of the 4-pt function leads to a recursion relation (eq.(21) and (26) of [46]) for the structure constants, similar to the one (5.11) for the 2-pt functions, which can be solved to [17, 18]

$$C(\alpha_1, \alpha_2, \alpha_3) = F_b(\alpha_1, \alpha_2, \alpha_3 | R) \left(\pi \mu \gamma(b^2) b^{2-2b^2} \right)^{(Q-\sum_i \alpha_i)/b} e^{I_Q(\alpha_1, \alpha_2, \alpha_3)} , \quad (5.24)$$

where $F_\nu(\alpha_1, \alpha_2, \alpha_3 | R)$ is a two-parameter (ν and R) family periodic function of α_i s with period ν . It is symmetric under the permutations of α_i s. In the special case when $\alpha_3 \rightarrow 0$, it must reduce to $f_{\nu/2}(\alpha | R)$, where $\alpha = \alpha_1 = \alpha_2$, for the 2-pt function;⁸ $F_b(\alpha, \alpha, 0 | R) = f_{b/2}(\alpha | R)$. The function $I_Q(\alpha_1, \alpha_2, \alpha_3)$ is defined in the appendix.

A similar consideration with the (1, 2) degenerate field leads to

$$C(\alpha_1, \alpha_2, \alpha_3) = F_{1/b}(\alpha_1, \alpha_2, \alpha_3 | -R) \left(\pi \tilde{\mu} \gamma(1/b^2) b^{-2+2/b^2} \right)^{b(Q-\sum_i \alpha_i)} e^{I_Q(\alpha_1, \alpha_2, \alpha_3)} , \quad (5.25)$$

Two results (5.24) and (5.25) have to match up. It then requires that

$$F_b(\alpha_1, \alpha_2, \alpha_3 | R) = e^{R(Q-\sum_{i=1}^3 \alpha_i)} F_{1/b}(\alpha_1, \alpha_2, \alpha_3 | -R) , \quad (5.26)$$

similar to the one (5.15) for the 2-pt functions. Note that this indeed reduces to that for the 2-pt functions when α_3 is sent to zero.

5.2.1 The $b = i$ case

We now focus on the $b = i$ case. The (dual-)CC dependent factors in (5.24) and (5.25) are to be replaced by (5.6) and (5.8) respectively, as in the case of 2-pt functions. The nontrivial functional part $e^{I_Q(\alpha_1, \alpha_2, \alpha_3)}$ in this limit was first found by Runkel and Watts [14] as a $c = 1$ limit of the minimal models for the spacelike theory, and then obtained as a certain $c_L = 1$ limit of the Liouville theory by Schomerus and Fredenhagen [15, 16]. In particular the 3-pt functions for the timelike theory were proposed in [15]. Since the explicit form of this function is not relevant at the level of our discussion, we will leave it for their original papers.

Similarly to the constraint (5.19) from the \mathbb{Z}_2 -symmetry, we have

$$F_{-i}(\alpha_1, \alpha_2, \alpha_3 | -R) = F_i(-\alpha_1, -\alpha_2, -\alpha_3 | R) . \quad (5.27)$$

⁸When α_3 is sent to zero, the structure constant $C(\alpha_1, \alpha_2, \alpha_3)$ vanishes unless $\alpha_2 = \alpha_1$ or $\alpha_2 = Q - \alpha_1$. In the $b = i$ limit, however, this point is subtle and requires a slightly nontrivial procedure [14].

Thus the momentum restriction is given by

$$\frac{F_i(\alpha_1, \alpha_2, \alpha_3|R)}{F_i(-\alpha_1, -\alpha_2, -\alpha_3|R)} = e^{-2i \sum_{i=1}^3 \alpha_i \ln \lambda_{lft}} = (-1)^{\sum_{i=1}^3 n_i} . \quad (5.28)$$

where we have plugged $\alpha_i = \omega_i = \pi n_i / 2 \ln \lambda_{lft}$ in, in the last equality. So for the constraint (5.22) for the 2-pt functions to agree with the energy quantisation (3.10) and (4.29), the l.h.s. cannot be equal to -1 , since otherwise it would have implied that $f_{i/2}(\alpha|R) = -f_{i/2}(-\alpha|R)$. Hence the l.h.s. must be equal to 1 for the agreement of the CFT construction with the TShG and matrix model results. It is then required that

$$\sum_{i=1}^3 n_i = 2n , \quad n \in \mathbb{Z}_+ . \quad (5.29)$$

This may simply imply the selection rule for the sum of energies in the 3-pt correlations. As noted in the end of the last subsection, the choice $F_i(\alpha_1, \alpha_2, \alpha_3|R) = 1$ may be preferred.

In sum, it can be consistent at this stage to have the discrete spectrum

$$\omega = \frac{\pi n}{2 \ln \lambda_{lft}} = \frac{\pi n}{2 \ln (\lambda_{ren} / (1 - \lambda_{ren}^2)^2)} , \quad n \in \mathbb{Z}_+ , \quad (5.30)$$

in constructing the $c_L = 1$ (thus $b = i$) limit of the timelike continuation of (2.1). The TShG and matrix model results seem to suggest that the unspecified functions $f_{i/2}(\alpha|R)$ and $F_i(\alpha_1, \alpha_2, \alpha_3|R)$ may simply be equal to 1 (together with a field renormalisation $V_\omega \rightarrow (e^{\pi i \tilde{\mu}_{ren}} / \mu_{ren})^{i\omega/2} V_\omega$), as noted before.

However, the most important check is the compatibility of our proposed spectrum with the conformal bootstrap equations [47]. The conformal bootstrap might fix the possible form of the function $F_i(\alpha_1, \alpha_2, \alpha_3|R)$ (so $f_{i/2}(\alpha|R)$), or in the worst case disprove any form to be inconsistent. To test the conformal bootstrap is beyond the scope of our paper. It remains to be seen.

For the spacelike $c_L = 1$ Liouville theory without the dual Liouville potential (or depending on the interpretation, with the self-dual Liouville potential, that is the $R = 0$ case in our notation), the conformal bootstrap was checked numerically by Runkel and Watts [14], whereas for the timelike $c_L = 1$ Liouville theory of [15, 19], it is yet to be checked, since in particular the timelike theory cannot be obtained by a naive analytic continuation as emphasised in [15].

5.3 Limiting Cases

As noted already, the energy quantisation (5.30) mimics exactly that of the system in a finite time interval ΔT , where

$$\Delta T = 2 \ln \lambda_{lft} = 2 \ln \left[\frac{\lambda_{ren}}{(1 - \lambda_{ren}^2)^2} \right] . \quad (5.31)$$

In fact the interval $|\Delta T|$ is the distance between two (renormalised) Liouville walls.

Here we would like to comment on two extreme cases, (1) $\Delta T \rightarrow \infty$ and (2) $\Delta T \rightarrow 0$. We are interested in the limiting spectrum and the (dual-)CC dependent factor in the 2 and 3-pt functions. We will restrict ourselves to the range $0 \leq \lambda_{ren} \leq 1$, since the renormalisation of the couplings (5.6) and (5.8) were obtained in the form of geometric series with the convergence radius $|\lambda_{ren}| < 1$ as shown in the appendix.

In the case (1), there are two ways to reach the limit; $\lambda_{ren} \rightarrow 0$ and $\lambda_{ren} \rightarrow 1$. We take the $\lambda_{ren} \rightarrow 0$ limit by

$$\tilde{\mu}_{ren} \rightarrow 0, \quad n \rightarrow \infty, \quad \text{keeping } \mu_{ren} \text{ and } \omega \sim \frac{\pi n}{2 \ln \lambda_{ren}} \text{ fixed finite}, \quad (5.32)$$

$$\mu_{ren} \rightarrow 0, \quad n \rightarrow \infty, \quad \text{keeping } \tilde{\mu}_{ren} \text{ and } \omega \sim \frac{\pi n}{2 \ln \lambda_{ren}} \text{ fixed finite}. \quad (5.33)$$

In both cases the spectrum becomes continuous, as the interval $|\Delta T|$ is now infinite. The (dual-)CC dependent factor for each case simply takes the form respectively of

$$(i\pi\mu_{ren})^{i\sum_{i=1}^3 \omega_i}, \quad \text{and} \quad (-i\pi\tilde{\mu}_{ren})^{-i\sum_{i=1}^3 \omega_i}. \quad (5.34)$$

This is indeed the result one would expect for the $c_L = 1$ limit [19, 15] of the timelike Liouville theory with a single Liouville potential.⁹ These limits provide the ingredient Liouville theories in the TShG model in section 3. It also suggests the renormalisation of the TShG coupling

$$(\lambda_{TShG})_{ren} = \lambda_{ren} \quad \longrightarrow \quad \lambda_{left} = \frac{\lambda_{ren}}{(1 - \lambda_{ren}^2)^2}, \quad (5.35)$$

away from $|(\lambda_{TShG})_{ren}| \ll 1$.

We next take the $\lambda_{ren} \rightarrow 1$ limit (thus $\pi\mu_{ren} = 1/\pi\tilde{\mu}_{ren}$) by

$$\lambda_{ren} \rightarrow 1, \quad n \rightarrow \infty, \quad \text{keeping } \omega \sim -\frac{\pi n}{4 \ln(1 - \lambda_{ren}^2)} \text{ fixed finite}. \quad (5.36)$$

Again the spectrum becomes continuous for the same reason, and the CC dependent factor reduces to

$$(-1)^{\frac{1}{2}\sum_{i=1}^3 n_i} \times (i\pi\mu_{ren})^{i\sum_{i=1}^3 \omega_i}. \quad (5.37)$$

The sign factor $(-1)^{\frac{1}{2}\sum_{i=1}^3 n_i}$ may be absorbed into a field renormalisation. Thus this case is essentially equivalent to the $\lambda_{ren} \rightarrow 0$ cases.

In the case (2), we need to take the coupling λ_{ren} to the critical value

$$\lambda_c = \{\lambda \mid \lambda^2 + \sqrt{\lambda} - 1 = 0; 0 \leq \lambda \leq 1\}. \quad (5.38)$$

This corresponds to $\lambda_{MM} = 1$ of the matrix model. The spectrum disappears, as the interval $|\Delta T|$ is now zero. As remarked before, this point is an analogue of $\lambda_o = \frac{1}{2}$ for the full S-brane of D-brane decay.

However, we note that since $\Delta T = 0$ is nothing but the self-dual point $R = 0$, we have an alternative isolated solution to the consistency conditions (5.20) and (5.28) – the continuous spectrum, which thus is essentially equivalent to the $c_L = 1$ limit of [19, 15].

⁹However, see the comment made right before the section 5.1.1.

6 Summary and discussions

We have proposed that the closed string energy in the bulk bouncing tachyon background is to be quantised as if strings were trapped in a finite time interval. The massless limit of the TShG model provides us with the most intuitive argument for why and how the energy should be quantised. The dual matrix model of the $(1+1)$ -dimensional string theory helps us to take the common speculation on the bulk tachyon condensation more seriously, and make the analogy between the bulk bouncing tachyon and the full S-brane of D-brane decay clearer. It has then proved to give a strong support for our proposal, manifesting the intuition of strings in a finite time interval in a concise manner. Our attempt of constructing the $c_L = 1$ limit of the Liouville theory with the dual Liouville potential has ended short of completion. It, however, has provided a different way to reach the energy quantisation from the CFT viewpoint, which can be consistent with the results in the TShG and matrix models, at least at the level of our analysis. The crucial test is to check whether our proposed discrete energy spectrum is compatible with the conformal bootstrap constraints. This remains to be seen.

The bulk tachyon might not have as interesting applications nor as much motivation to consider it as the open string tachyon. However, the matrix cosmology of the bulk bouncing tachyon offers an interesting toy model of the Big-Bang/Crunch cosmology. It might be interesting to consider an application of the bulk bouncing tachyon to cosmology in higher dimensions, though in which taking into account the back-reaction from the string pair productions to the geometry would be a challenging problem.

Finally there is a certain parallel between the bulk bouncing tachyon and the Schwarzschild bubble of [51]. It might be worthwhile to revisit the question discussed in [52].

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Appendix A: Calculation of Structure Constant

In this appendix, we will show the details of the computation of the structure constant $C_{\alpha, -b/2}^{\alpha+b/2}$, in particular, with emphasis on the peculiarity of the $b = i$ case.

From the OPE (5.2), the structure constant may be given by

$$C_{\alpha, -b/2}^{\alpha+b/2} = \lim_{\substack{z_1 \rightarrow \infty, z_3 \rightarrow 0 \\ z_2 \rightarrow 1}} |z_1|^{4\Delta_Q - \alpha - b/2} \sum_{n,m=0}^{\infty} \frac{(-\mu)^n (-\tilde{\mu})^m}{n!m!} \quad (A.1)$$

$$\times \langle V_{Q-\alpha-b/2}(z_1) V_{\alpha}(z_2) \left(\int d^2u e^{2b\phi(u)} \right)^n \left(\int d^2v e^{2\phi(v)/b} \right)^m V_{-b/2}(z_3) \rangle_{free} ,$$

where $\langle \cdots \rangle_{free}$ denotes the correlation function of free CFT. Except for $b = i$, the only contribution comes from $(n, m) = (1, 0)$ in the sum, due to the “charge” conservation in the presence of the background charge $-2Q$. Then we only need to evaluate the single integral, which gives us (5.4).

However, in the case of $b = i$, since the combination $\int d^2 z e^{2b\phi(z)} \int d^2 w e^{2\phi(w)/b}$ of two screening integrals has zero net charge, there is an infinite number of contributions from all the terms of $(n, m) = (k+1, k)$ with $k \in \mathbb{Z}_+ \cap \{0\}$. Therefore we now need to evaluate the Dotsenko-Fateev integral [53]

$$I_{m,n}(\{\alpha_i\}; b) \equiv \lim_{\substack{z_1 \rightarrow \infty, z_3 \rightarrow 0 \\ z_2 \rightarrow 1}} |z_1|^{4\Delta_{\alpha_1}} \frac{(-\mu)^n (-\tilde{\mu})^m}{n!m!} \\ \times \int \langle \prod_{i=1}^3 V_{\alpha_i}(z_i) \prod_{k=1}^n V_b(u_k) \prod_{j=1}^m V_{1/b}(v_j) \rangle_{free} d^2 u_k d^2 v_j, \quad (\text{A.2})$$

with $(n, m) = (k+1, k)$ with $k \in \mathbb{Z}_+ \cap \{0\}$.

This integral can be read off from the Liouville 3-pt function, since the 3-pt function was so constructed that it obeys

$$I_{n,m}(\alpha_i; b) = \text{Res}_{\sum_i \alpha_i = Q - nb - m/b} C(\alpha_1, \alpha_2, \alpha_3), \quad (\text{A.3})$$

where the 3-pt function is given by

$$C(\alpha_1, \alpha_2, \alpha_3) = \left(\pi \mu \gamma(b^2) b^{2-2b^2} \right)^{(Q - \sum_i \alpha_i)/b} e^{I_Q(\alpha_1, \alpha_2, \alpha_3)}, \quad (\text{A.4})$$

and we have defined

$$e^{I_Q(\alpha_1, \alpha_2, \alpha_3)} \equiv \frac{\Upsilon_0 \Upsilon(2\alpha_1) \Upsilon(2\alpha_2) \Upsilon(2\alpha_3)}{\Upsilon(\sum_i \alpha_i - Q) \Upsilon(\alpha_1 + \alpha_2 - \alpha_3) \Upsilon(\alpha_2 + \alpha_3 - \alpha_1) \Upsilon(\alpha_3 + \alpha_1 - \alpha_2)}, \quad (\text{A.5})$$

and the replacement

$$\left[\pi \mu \gamma(b^2) b^{2-2b^2} \right]^{(Q - \sum_i \alpha_i)/b} \Big|_{\sum_i \alpha_i = Q - nb - m/b} \rightarrow \left[\pi \mu \gamma(b^2) b^{2-2b^2} \right]^n \left[\pi \tilde{\mu} \gamma(1/b^2) (1/b)^{2-2/b^2} \right]^m,$$

is required in (A.4). The special function $\Upsilon(x)$ is defined by

$$\ln \Upsilon(x) \equiv \int_0^\infty \frac{dt}{t} \left[(Q/2 - x)^2 e^{-2t} - \frac{\sinh^2(Q/2 - x)t}{\sinh(bt) \sinh(t/b)} \right], \quad (\text{A.6})$$

and

$$\Upsilon_0 \equiv \left. \frac{d\Upsilon(x)}{dx} \right|_{x=0}. \quad (\text{A.7})$$

It has zeroes at $x = -nb - m/b$ and $x = Q + nb + m/b$ with $n, m \in \mathbb{Z}_+ \cap \{0\}$. Also it is clearly invariant under $b \leftrightarrow 1/b$.

By making use of the formula

$$\begin{aligned} \Upsilon(x - nb - m/b) &= (-1)^{nm} b^{-n(n+1)b^2} b^{m(m+1)/b^2} (b^{-1+2bx})^n (b^{1-2x/b})^m \\ &\times \left[\prod_{l=1}^n \prod_{k=1}^m (x - lb - k/b)^{-2} \prod_{l'=1}^n \gamma(bx - l'b^2) \prod_{k'=1}^m \gamma(x/b - k'/b^2) \right]^{-1} \Upsilon(x), \end{aligned} \quad (\text{A.8})$$

one can find that

$$\begin{aligned} I_{n,m}(\{\alpha_i\}; b) &= b^{4(n-m)} [\pi\mu\gamma(b^2)]^n [\pi\tilde{\mu}\gamma(1/b^2)]^m \\ &\times \prod_{l=1}^n \prod_{k=1}^m (lb + k/b)^{-2} \prod_{l'=1}^n \gamma(-l'b^2) \prod_{k'=1}^m \gamma(-k'/b^2) \\ &\times \prod_{i=1}^3 \left[\prod_{l''=0}^{n-1} \prod_{k''=0}^{m-1} (2\alpha_i + l''b + k''/b)^{-2} \prod_{l'''=0}^{n-1} \gamma(-2b\alpha_i + 1 - l'''b^2) \prod_{k'''=0}^{m-1} \gamma(-2\alpha_i/b + 1 - k'''/b^2) \right]. \end{aligned} \quad (\text{A.9})$$

We now apply this formula to the $b = i$ case by taking the following limit

$$C_{\alpha, -i/2}^{\alpha+i/2} = \lim_{\epsilon \rightarrow 0_+} \lim_{\eta \rightarrow 0} \sum_{n=0}^{\infty} I_{n+1,n}(Q - \alpha - b/2, \alpha, -b/2 + \eta; b) \Big|_{b^2 = -1+i\epsilon}, \quad (\text{A.10})$$

where we first take the $\eta \rightarrow 0$ limit and then $\epsilon \rightarrow 0_+$. Without this limiting procedure, the Dotsenko-Fateev integral (A.9) is not well-defined. Note that the result crucially depends on how to perform the regularisation and take the limit. Our reasoning for choosing this particular limit is the following; There is a zero and pole at $\alpha = -b/2$ in (A.9) irrespective of the value of b , so we first regularise them by introducing the regulator η for arbitrary b . Then since there appear additional zeroes and poles when $b^2 \rightarrow -1$, we further introduce the regulator ϵ . A particular choice $b^2 = -1 + i\epsilon$ is instructed from Schomerus's construction of the $c_L = 1$ Liouville 3-pt functions [15].

Using repeatedly the formulae

$$\begin{aligned} \gamma(x - n) &= (-1)^n \prod_{l=1}^n (x - l)^{-2} \gamma(x), & \gamma(x + n) &= (-1)^n \prod_{l=0}^{n-1} (x + l)^2 \gamma(x), \\ \gamma(x) \gamma(1 - x) &= 1, & \gamma(x) \gamma(-x) &= -\frac{1}{x^2}, \end{aligned} \quad (\text{A.11})$$

one can find that

$$C_{\alpha, -i/2}^{\alpha+i/2} = i\pi\mu_{ren} \sum_{n=0}^{\infty} (n+1) (\pi\mu_{ren})^n (\pi\tilde{\mu}_{ren})^n = \frac{-i\pi\mu_{ren}}{(1 - (\pi\mu_{ren})(\pi\tilde{\mu}_{ren}))^2}, \quad (\text{A.12})$$

where we have renormalized the CCs by taking the limits $\mu, \tilde{\mu} \rightarrow 0$ and $\epsilon \rightarrow 0_+$ with $b^2 = -1 + i\epsilon$, while $\mu_{ren} \equiv -i\mu\gamma(b^2)$ and $\tilde{\mu}_{ren} \equiv i\tilde{\mu}\gamma(b^2)$ kept fixed, as in Section 3. Since the series in (A.12) is only convergent for $\lambda_{ren}^2 = (\pi\mu_{ren})(\pi\tilde{\mu}_{ren}) < 1$, either the theory with $|\lambda_{ren}| > 1$ is ill-defined, or one might take the final expression of (A.12) to define the theory with $|\lambda_{ren}| > 1$ by means of analytic continuation.

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